

RUBIN OBSERVATORY LEGACY SURVEY OF SPACE AND TIME (LSST)

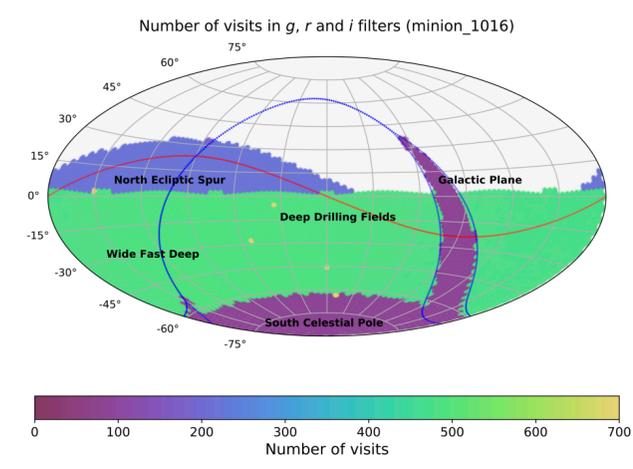


Effective Mirror Diameter	6.7 m
Field of view	9.6 sq deg
Survey length	10 years
Sky coverage	~18,000 sq deg
Site	Cerro Pachon
Filters	ugrizy
Typical seeing	0.7"
Exposure ('Visit') Time	2x15 s /visit
Readout Time	2 s
Typical slew/settle	5 s
Data rate	~15 TB/night
Photometric accuracy	10 mmag
Astrometric accuracy	50 mas
Astrometric precision	10 mas

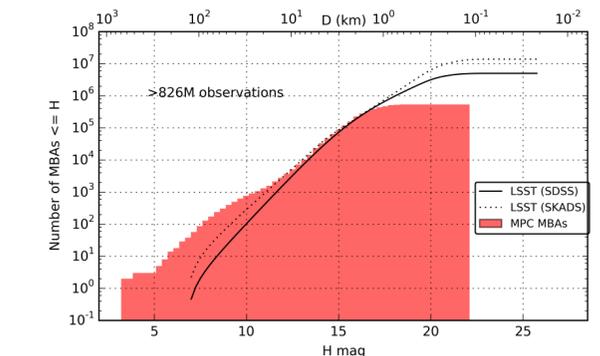
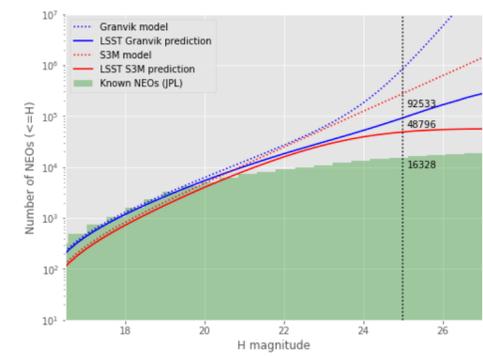
	u	g	r	i	z	y
Typical # visits/field	70	100	230	230	200	200
m5 depth	23.8	24.8	24.4	24.0	23.3	22.4



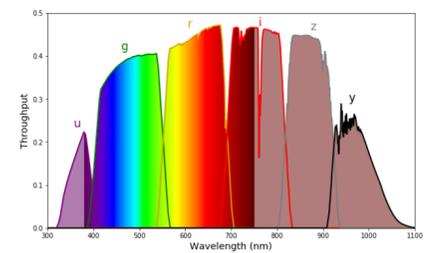
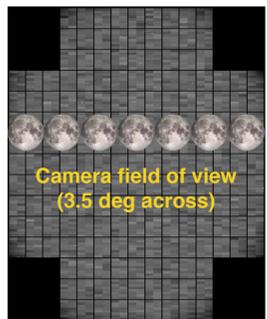
	Currently Known*	LSST Discoveries**	Typical number of observations+
Near Earth Objects (NEOs)	~25,500	100,000	(D>250m) 60
Main Belt Asteroids (MBAs)	~1,000,000	5,000,000	(D>500m) 200
Jupiter Trojans	~10,000	280,000	(D>2km) 300
TransNeptunian Objects (TNOs) + Scattered Disk Objects (SDOs)	~4000	40,000	(D>200km) 450
Comets	~4000	10,000	?
Interstellar Objects (ISOs)	2	>10	?



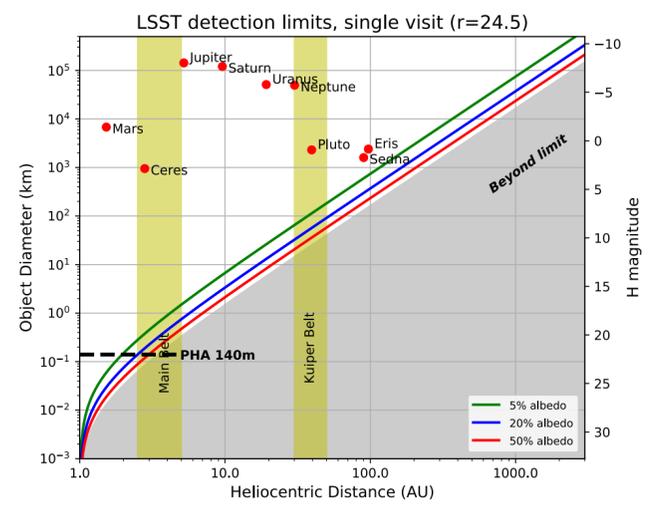
Survey strategy details are evolving; generally, each field observed 2x / night, every 3 nights. Footprint and details of cadence under optimization.



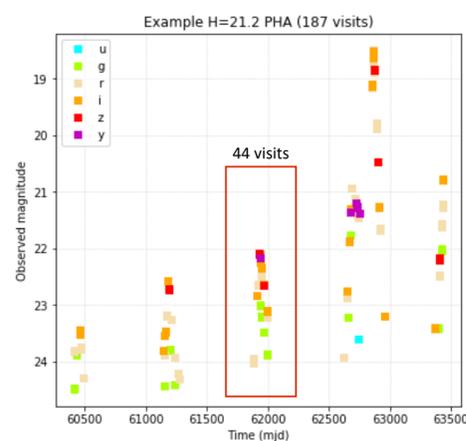
The uncertainty in the size distribution for most solar system populations leads to fairly large uncertainties in the total number of objects the Rubin Observatory is expected to discover after 10 years, however the survey is expected to detect on the order of 100,000 NEOs and between 5 to 9 million MBAs.



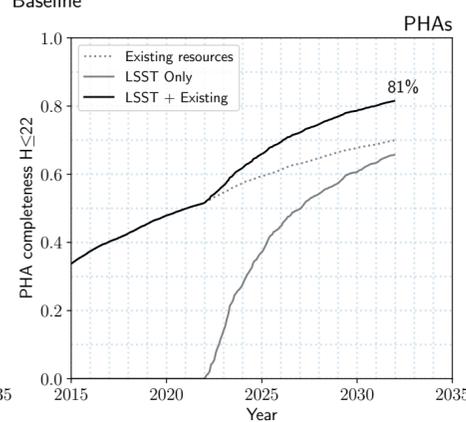
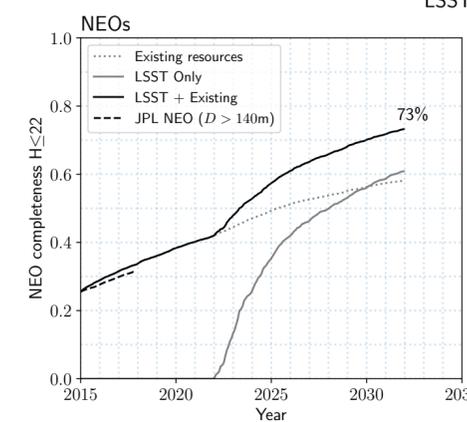
Throughput curves



Sensitivity limits (single visit)



Simulated observations of a PHA with H=21.



Jones et al, 2018