

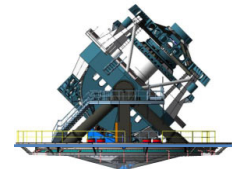


# Solar System Metrics & Survey Strategy

Lynne Jones & Peter Yoachim

LSST2019 Project & Community Workshop  
August 13, 2019

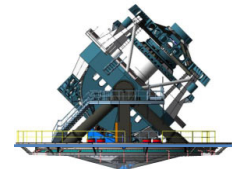




## Current Status



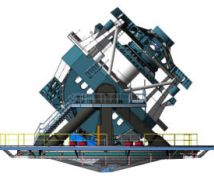
- WP Nov 2018 / SAC guidelines May 2019 / Simulations..
- Investigation of many families ..
  - See July2019 update on community with more information on FBS 1.2 set of runs
  - <https://community.lsst.org/t/july-2019-update/3760>  
(<http://ls.st/xsb>)
- Not complete, but getting better understanding of what metrics are telling us and what will need to go into report to the SCOC
  - Working with subsets of the community, but looking to add more community posts surrounding this
- Will be releasing more runs and also writeups describing interpretation and comparisons of various runs



# Solar system metrics



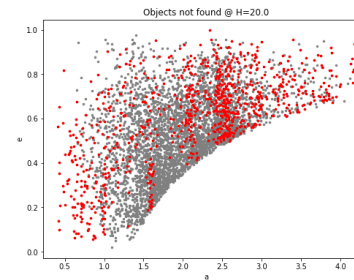
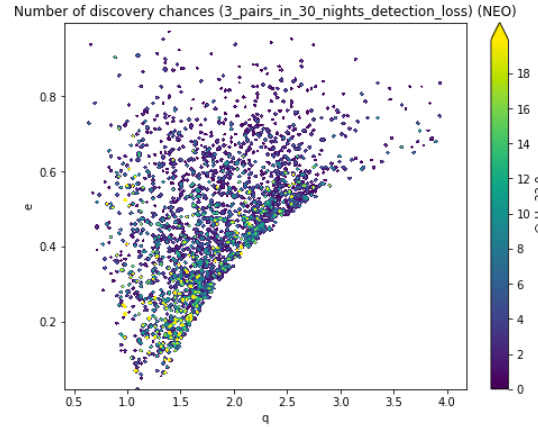
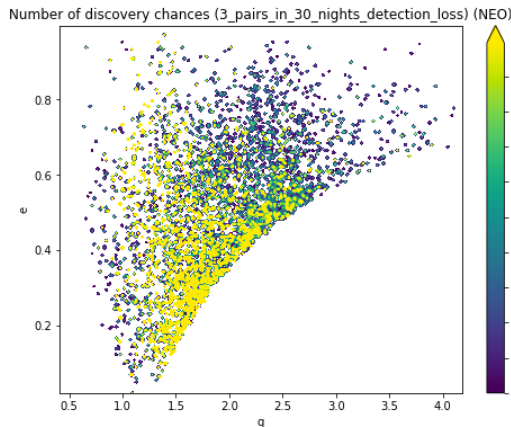
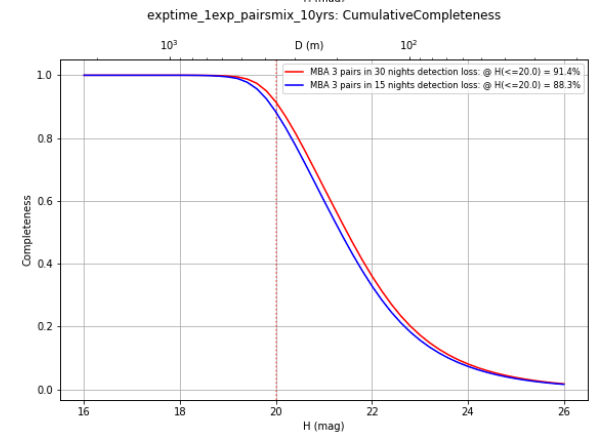
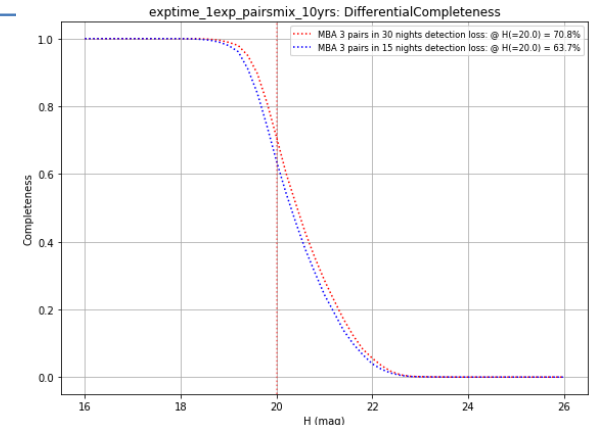
- To calculate metrics, first need to generate simulated observations of the objects
  - NEO (Granvik), MBA (S3M), Trojan (S3M), TNO (CFEPS L7)
  - + sims\_movingObjects
  - [https://github.com/lsst-sssc/SSSC\\_test\\_populations\\_gitlfs](https://github.com/lsst-sssc/SSSC_test_populations_gitlfs)

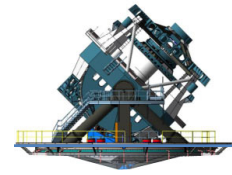


# Solar system metrics



- Discovery metrics (as previously)
  - 3 pairs in 15 nights
  - 3 pairs in 30 nights
  - .. 3 pairs in 12 nights, 20 nights
  - .. 4 pairs in 20 nights, 3 triplets in 30
  - .. Single detection, single pair
  - .. 3x15 & 3x30 @ SNR=3,4,5.
  - HighVelocity (trailing) pair



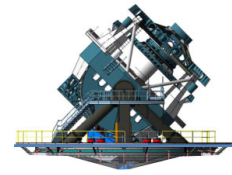


# Solar system metrics

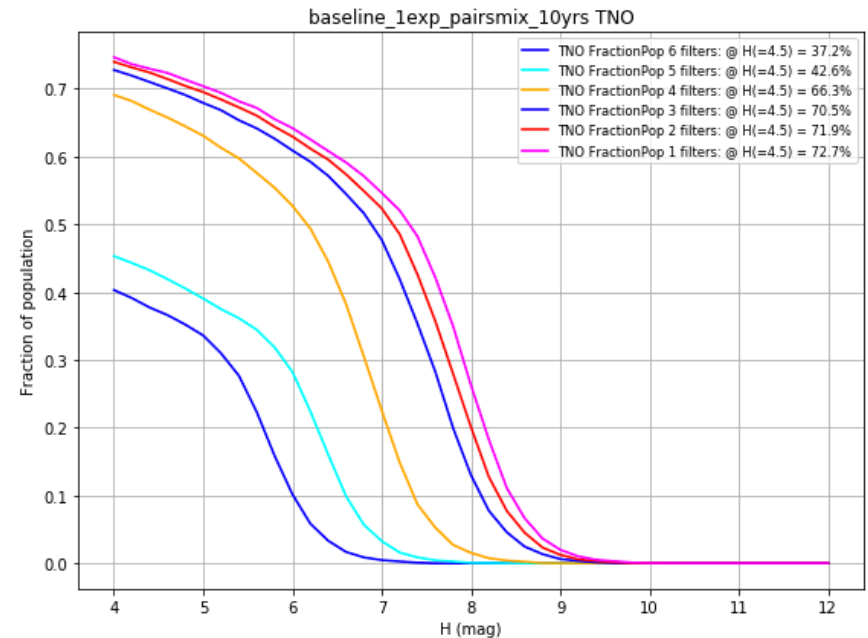
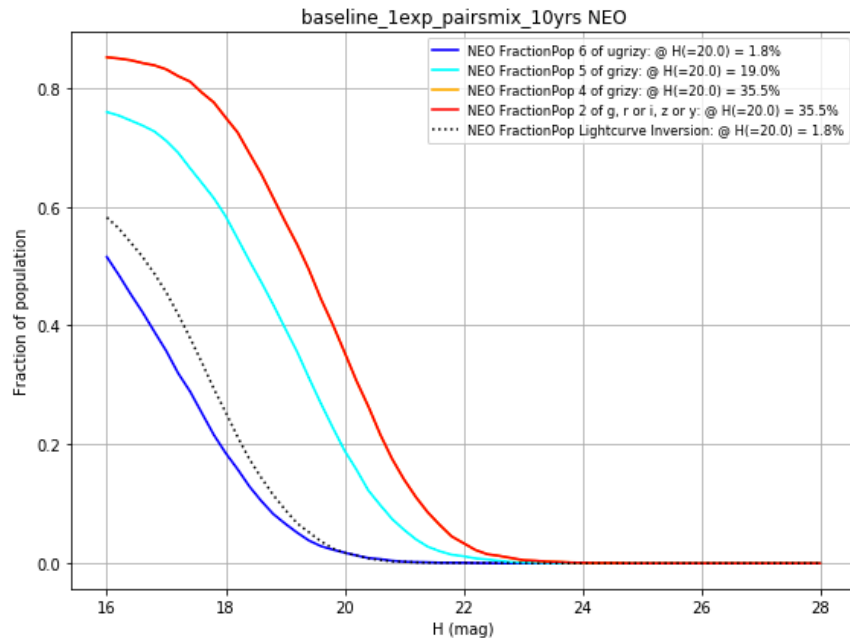


- Characterization metrics (as previously)
  - “Chance of detecting activity”
    - Bin time (or mean anomaly) over survey (or orbit) and build histogram of visits - what fraction of bins received a visit?
  - Inner solar system
    - Lightcurve inversion (updated metric)
    - Color determination (10+ SNR-weighted observations)
      - $g + ([r \text{ or } i] \text{ OR } [z \text{ or } y])$  (2 colors)
      - 4 of grizy
      - 5 of grizy
      - 6 of ugrizy
  - Outer solar system
    - Lightcurve/Color in 1, 2, 3, 4, 5 or 6 filters (30+ obs in first filter, 20+ in secondary)

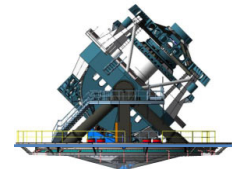




# Solar system metrics



For the runs described next, these plots for NEO, MBA, Trojan & TNO populations online at <http://astro-lsst-01.astro.washington.edu:8081> .. will also find a link for downloading full outputs if interest



# Solar system metrics



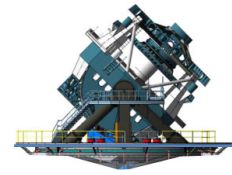
Pending:

Incorporate resonant TNO populations

\* Should get very red SED [Wes Fraser]

Incorporate MBC population

\* Need MBC magnitude and/or updated metric [Mike Kelly]



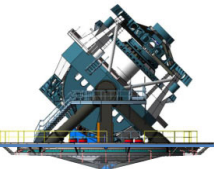
## Runs matching SAC requests

- Pairs in same or different filters
- 1x30 or 2x15 visits
- Presto color (g+i+pause+g)
- Footprint variations (9 different ones)
- Rolling cadences (13 variations)
- Dithered DDF (spatial)
- DESC DDF
- Target of opportunity (ToO)
- Vary u-band filter loading
- 1s or 5s exposure sky coverage
- Stability tests

## Bonus Experiments

- Pathological footprint
- Variable exposure time
- AltSched like behavior
- Camera rotator dithering
- Smarter rolling cadence



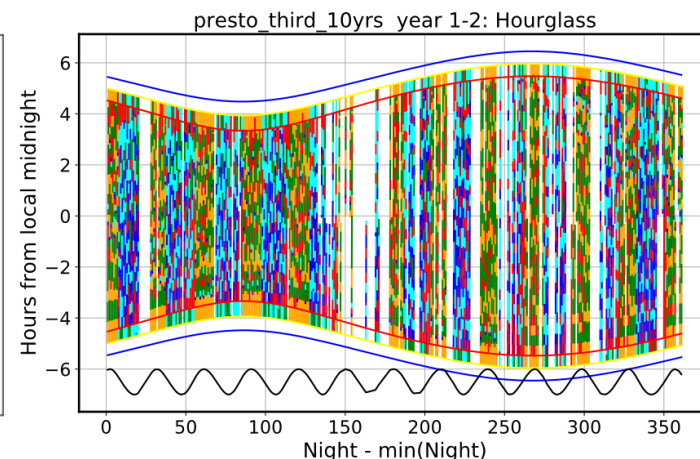
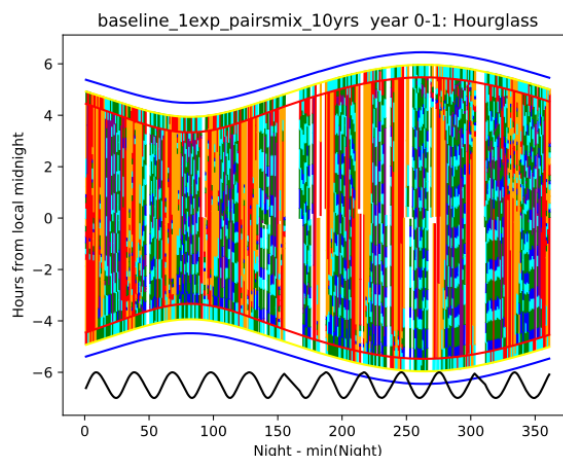
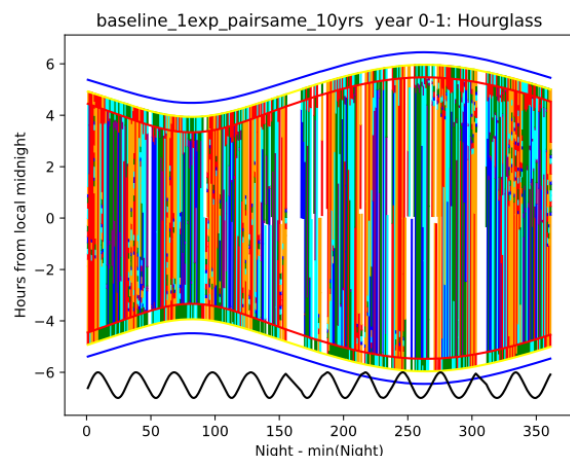


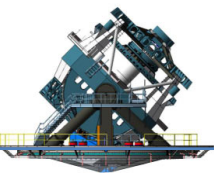
# Intra-night visits FBS 1.2 runs



## Baseline-like sims

- 1x30s snaps in a visit
- 2x15s snaps in a visit
- g+g, r+r, i+i pairs (u, z, y not paired) PairSame
- g+r, r+i, i+z pairs (u, some z, y not paired) PairMix
- g+i...40-120min later+g, r+z...+r Presto

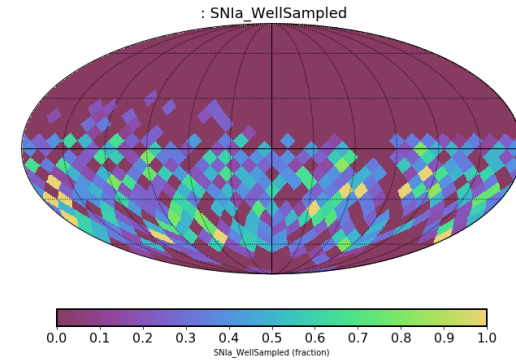
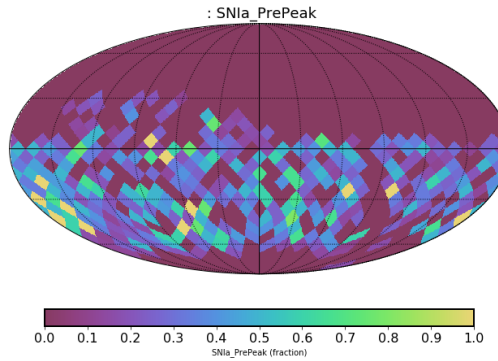
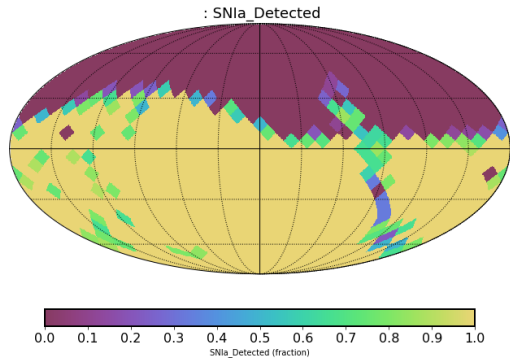




# Intra-night visits FBS 1.2 runs



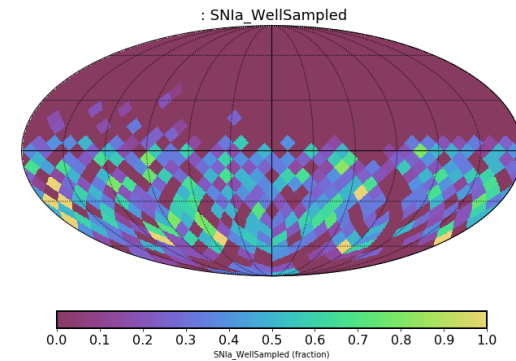
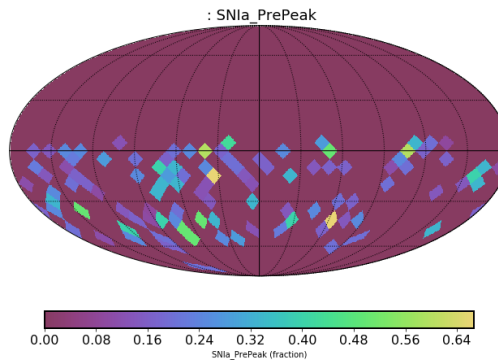
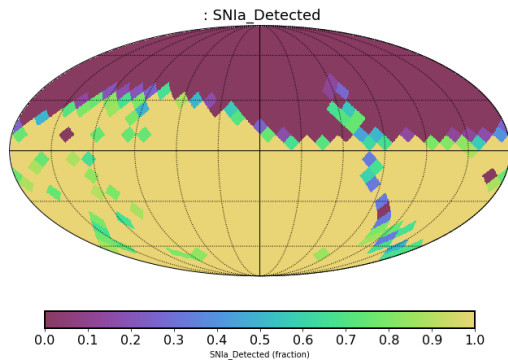
## 4k Type Ia SNe



Pairs mixed 62% detected

15% measured pre-peak

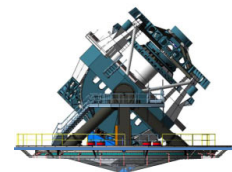
17% “well-sampled”



Pairs same 62% detected

3% measured pre-peak

16% “well-sampled”

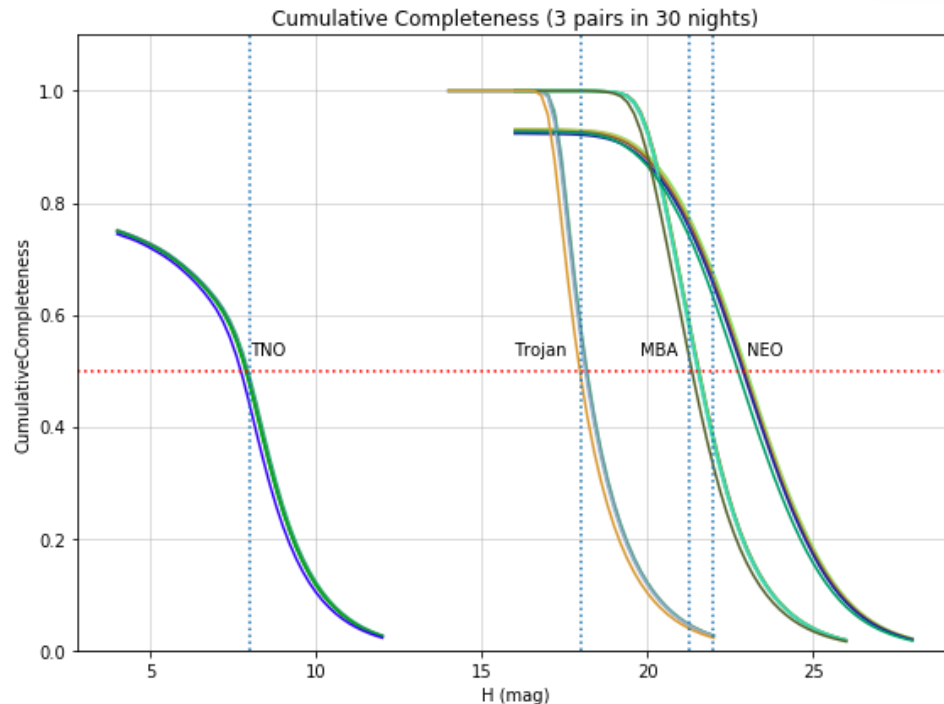


# Intra-night visits FBS 1.2 runs



Mixing filters for pairs does have a cost to SSOs.

Set up some small % of visits to be same filters?

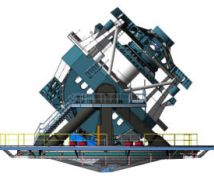


	NEO H=22	MBA H=21.25	Trojan H=18	TNO H=7.5
baseline_1exp_pairsame_10yrs	67.3	59.2	57.5	57.7
baseline_1exp_pairsmix_10yrs	66.5	58.7	56.6	57.0
baseline_2exp_pairsame_10yrs	66.0	58.2	56.3	56.9
baseline_2exp_pairsmix_10yrs	65.2	57.8	55.0	56.6
presto_third_10yrs	62.9	52.7	48.4	54.3

~1% pairs same -> mix

~2% mix pairs + 2x15s

~5-6% presto\_third

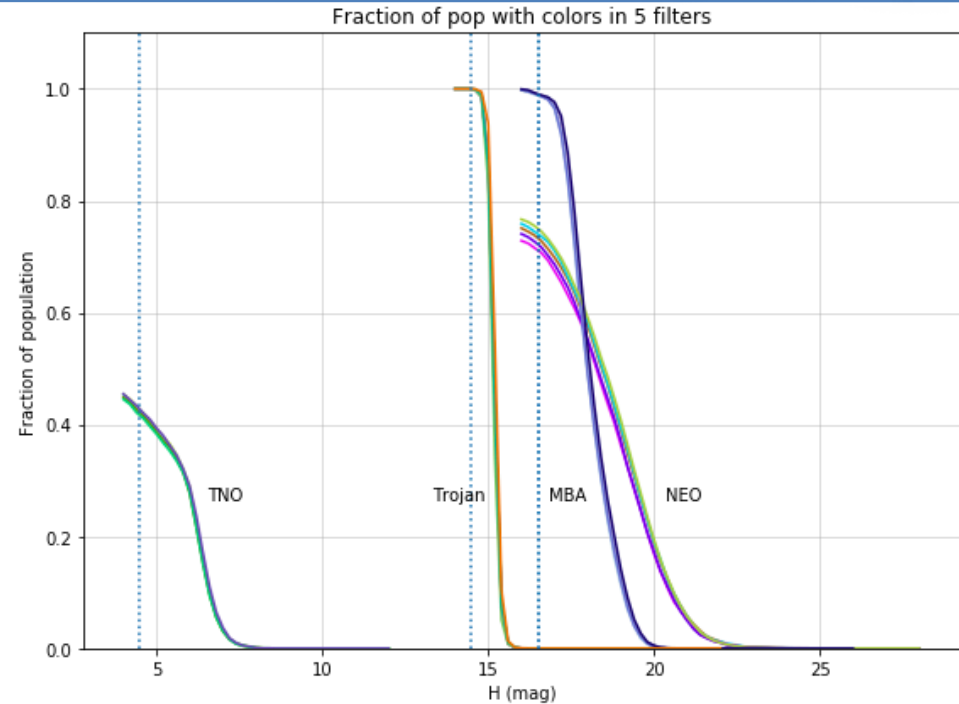


# Intra-night visits FBS 1.2 runs



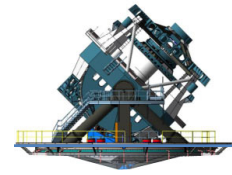
Characterization fairly insensitive to filters used for pairs\*.

\* presto\_third increases NEO characterization, but has a larger cost to overall discovery



**NEO H=16.5    MBA H=16.5    Trojan H=14.5    TNO H=4.5**

	NEO H=16.5	MBA H=16.5	Trojan H=14.5	TNO H=4.5
<b>baseline_1exp_pairsame_10yrs</b>	73.3	98.9	100.0	42.6
<b>baseline_1exp_pairsmix_10yrs</b>	74.1	99.0	100.0	42.6
<b>baseline_2exp_pairsame_10yrs</b>	71.1	98.8	100.0	42.1
<b>baseline_2exp_pairsmix_10yrs</b>	72.2	98.9	100.0	41.6
<b>presto_third_10yrs</b>	75.0	99.1	100.0	42.7



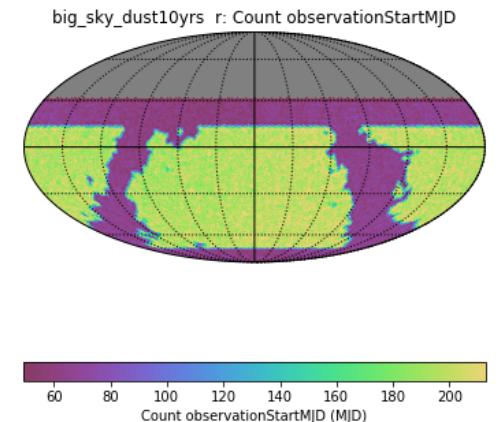
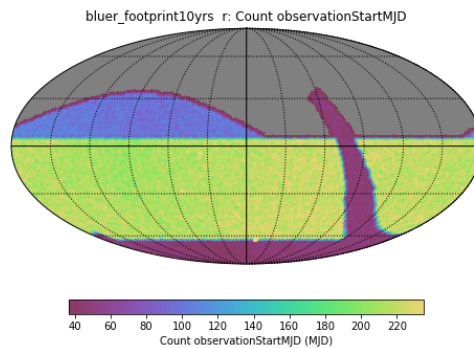
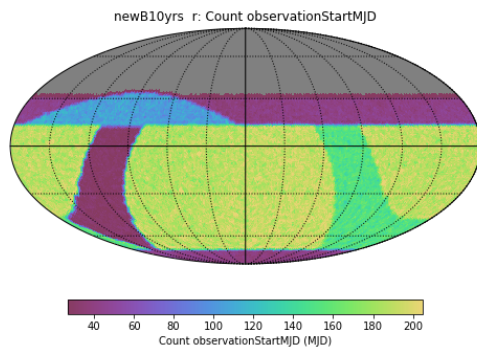
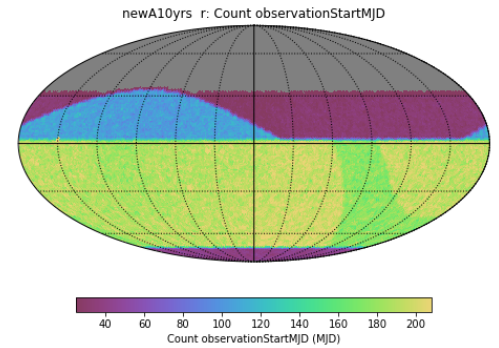
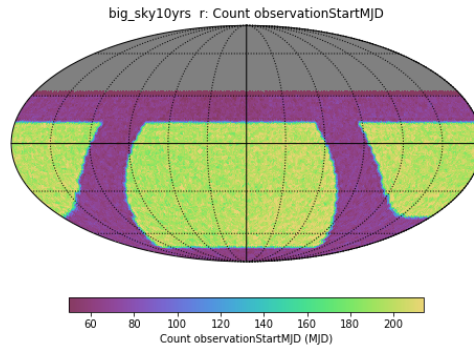
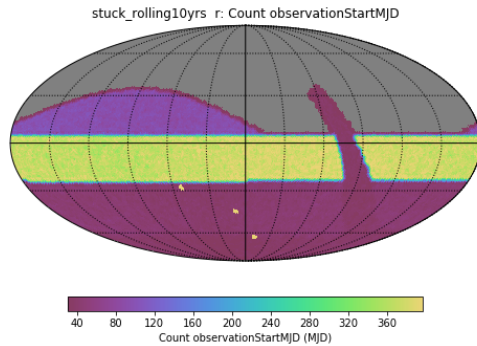
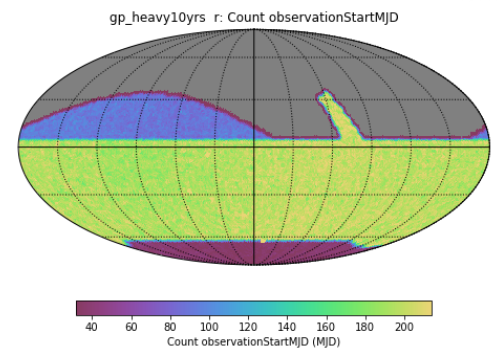
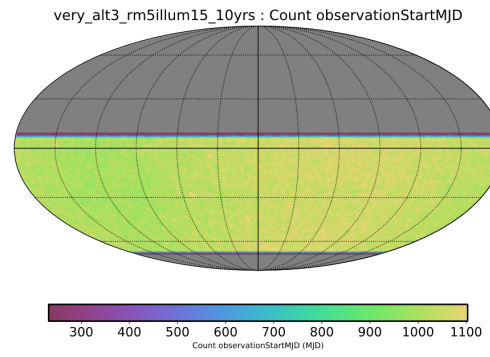
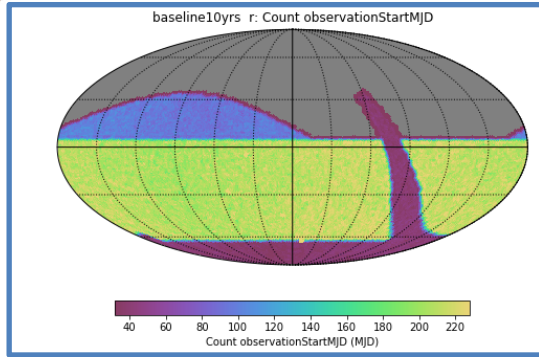
- Motivation for pairs in different filters comes from characterizing (slowly) changing transients
- Increases pre-peak sampling of SNIa (3% - 15%)
- Pushback comes from worry that SSOs will be harder to discover - there is some impact on SSO discovery, particularly with presto\_third in its current form.
- Additional pushback that changing filters is less efficient - ~2% penalty
- Add some small % of visits in same-filter [Volk wp]

	<b>Nvisits</b>
<b>baseline_1exp_nopairs_10yrs</b>	100.8
<b>baseline_1exp_pairsame_10yrs</b>	100.0
<b>baseline_1exp_pairsmix_10yrs</b>	97.9
<b>baseline_2exp_pairsame_10yrs</b>	92.4
<b>baseline_2exp_pairsmix_10yrs</b>	90.5

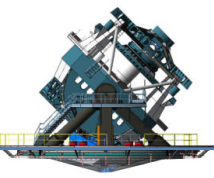




# FBS 1.2 runs : footprints





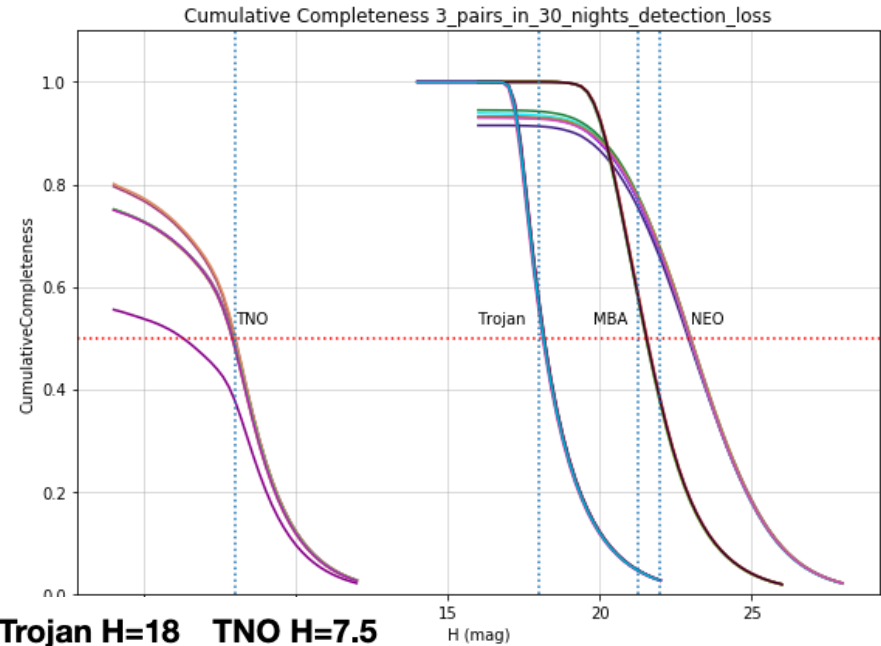


# WFD footprint FBS 1.2 runs

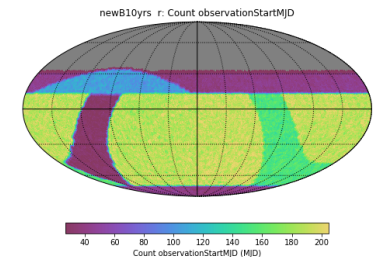


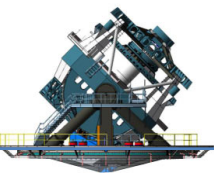
## - 'WFD footprint'

More coverage in the north improves discovery of TNOs.



	NEO H=22	MBA H=21.25	Trojan H=18	TNO H=7.5
baseline_1exp_pairsame_10yrs	67.3	59.2	57.5	57.7
baseline_1exp_pairesmix_10yrs	66.5	58.7	56.6	57.0
very_alt3_rm5illum50_10yrs	65.4	58.3	57.0	43.2
newA10yrs	67.0	58.0	54.9	59.5
newB10yrs	67.1	58.6	55.9	60.0
bluer_footprint10yrs	67.5	59.0	57.5	57.2
gp_heavy10yrs	66.8	58.5	56.1	57.1





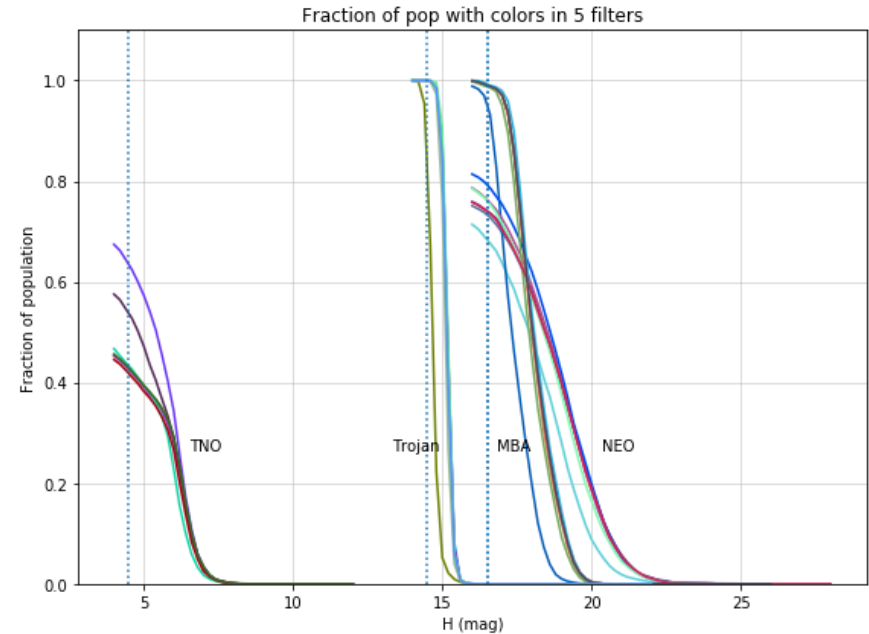
# WFD footprint FBS 1.2 runs



## - 'WFD footprint'

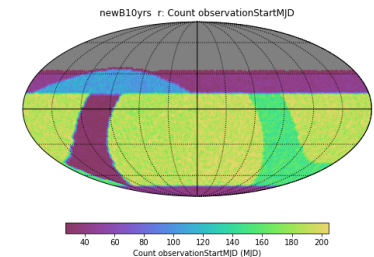
More coverage in the north improves characterization of TNOs AND NEOs.

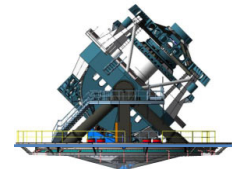
Need to add MBC & resonant TNO



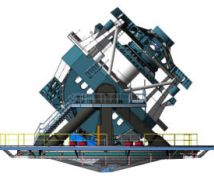
NEO H=22    MBA H=21.25    Trojan H=18    TNO H=7.5

	NEO H=22	MBA H=21.25	Trojan H=18	TNO H=7.5
<b>baseline_1exp_pairsame_10yrs</b>	67.3	59.2	57.5	57.7
<b>baseline_1exp_paiasmix_10yrs</b>	66.5	58.7	56.6	57.0
<b>very_alt3_rm5illum50_10yrs</b>	65.4	58.3	57.0	43.2
<b>newA10yrs</b>	67.0	58.0	54.9	59.5
<b>newB10yrs</b>	67.1	58.6	55.9	60.0
<b>bluer_footprint10yrs</b>	67.5	59.0	57.5	57.2
<b>gp_heavy10yrs</b>	66.8	58.5	56.1	57.1





- Footprint
- Motivated first by DESC desire for more galaxies
- Galaxy counts don't show clear improvement
- Some motivation from mini-surveys (NES, Euclid, DESI)
- Some improvement for SSOs (discovery and characterization) if WFD footprint extended north
  
- Return to DESC for clarification on metrics
- Redo footprint with  $E(B-V)$  cut exactly (done)
- Redo footprint with 90% WFD and with 825 visits .. look at options to increase visits in N/S (counter poor seeing)



# FBS 1.2 : rolling cadence

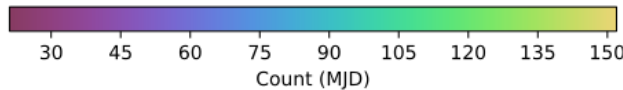
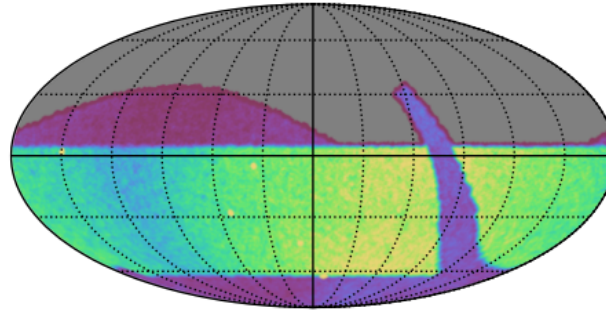


Start off normal, then divide WFD in (half) and alternate emphasis on north and south

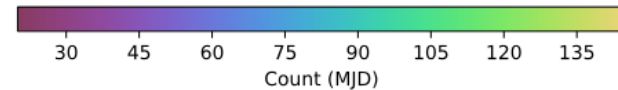
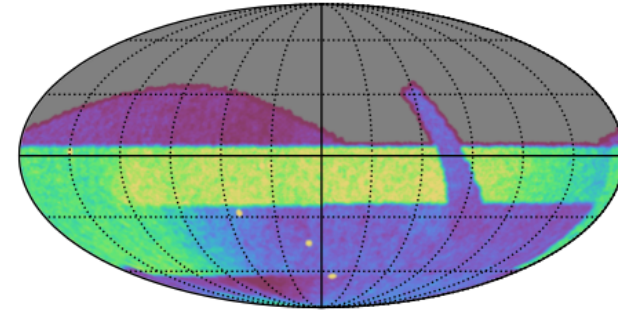
Year 1 like baseline, WFD gets 120 observations/yr

Rolling, get 25 or 215 observations per year

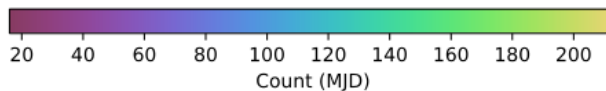
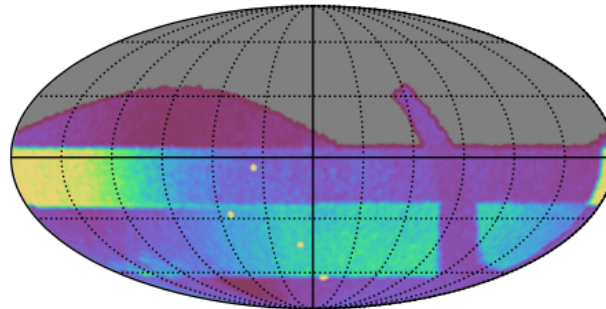
mod2 night > 0 and night < 365: Count



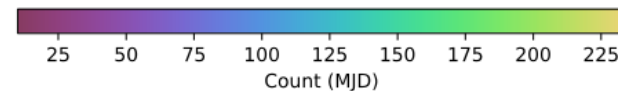
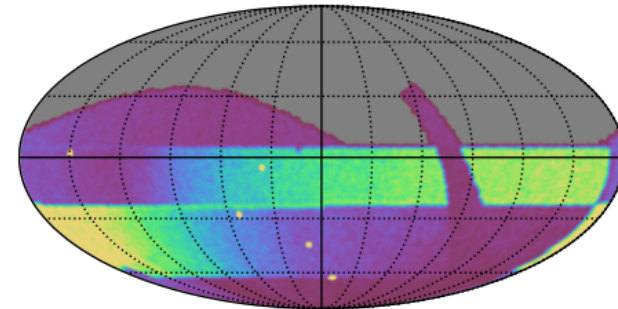
mod2 night > 365 and night < 730: Count

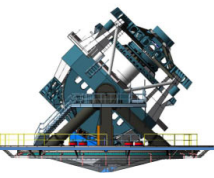


mod2 night > 730 and night < 1095: Count



mod2 night > 1095 and night < 1461: Count



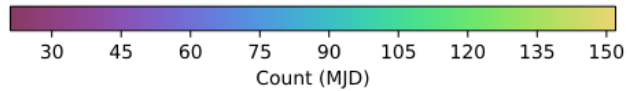
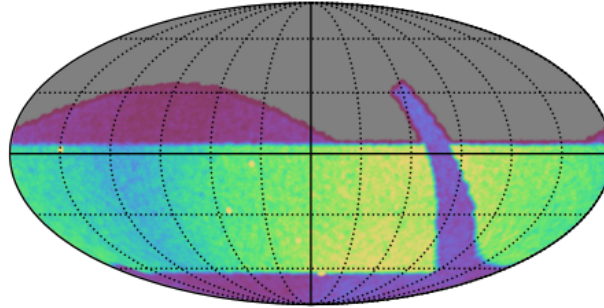


# FBS 1.2 : rolling cadence

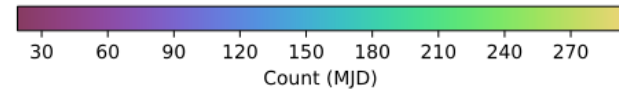
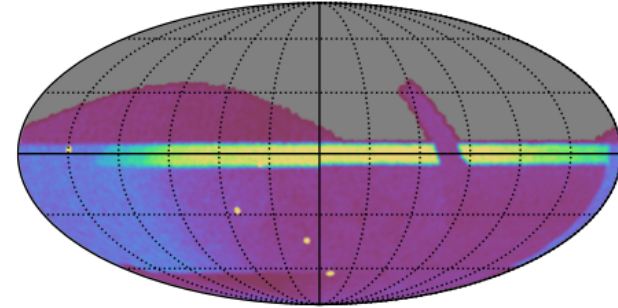


More extreme rolling: have 6 declination stripes.  
Then 450+ observations in a season, with 400 more visits over remaining 9 years.

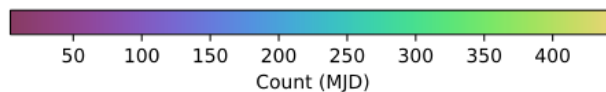
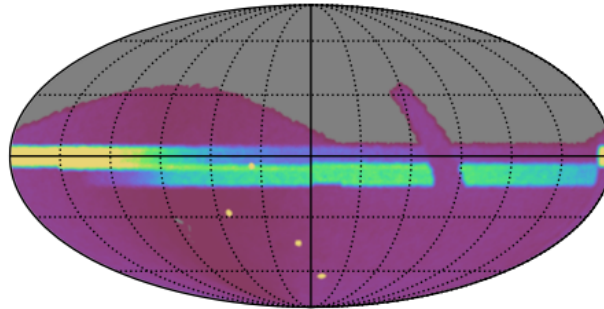
mod6 night > 0 and night < 365: Count



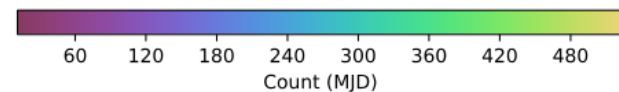
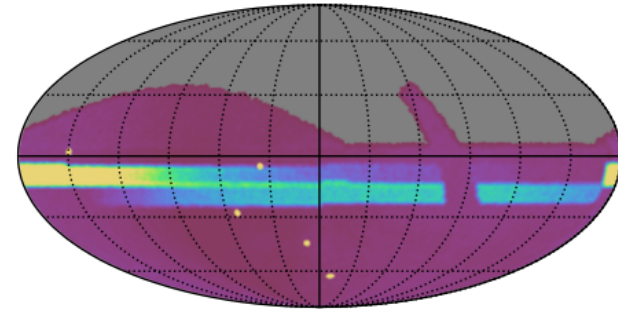
mod6 night > 365 and night < 730: Count



mod6 night > 730 and night < 1095: Count



mod6 night > 1095 and night < 1461: Count



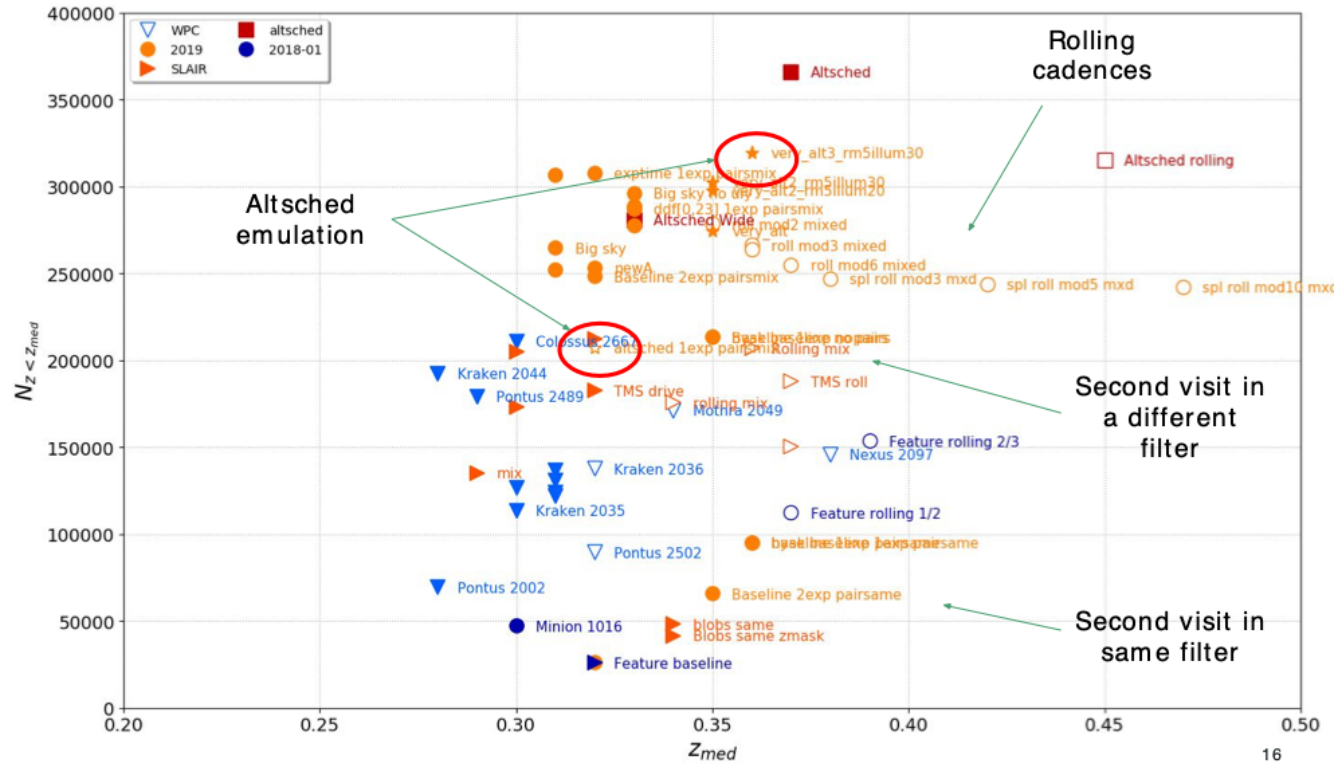




# Rolling cadence FBS 1.2 runs



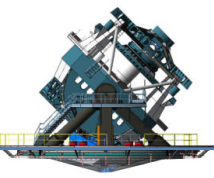
- ‘rolling cadence variations’



SNe group has been running more intensive analysis and giving feedback on sims

From Nicolas Regnault



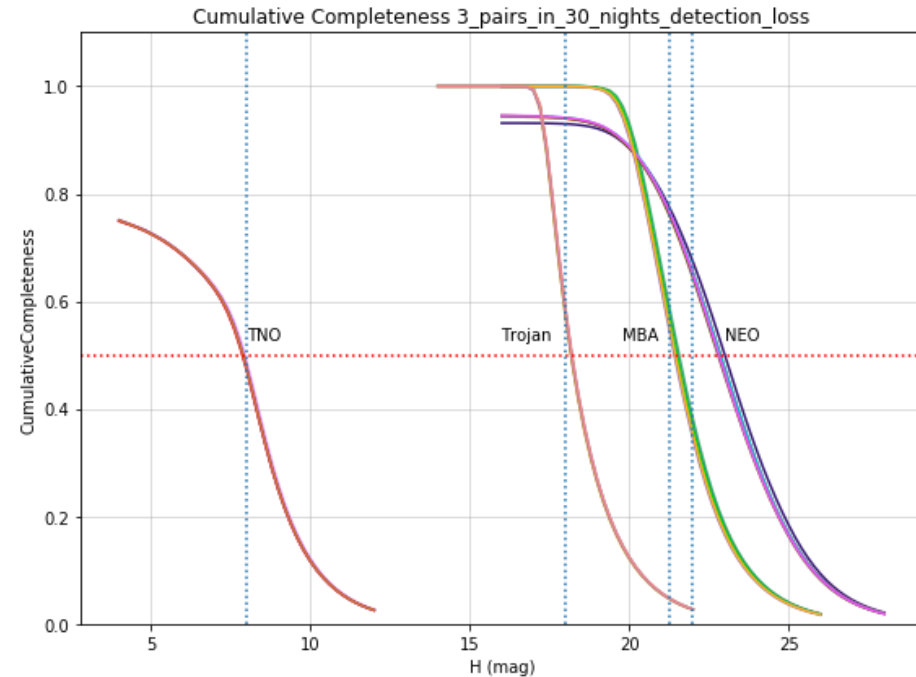


# Rolling cadence FBS 1.2 runs



– ‘rolling cadence variations’

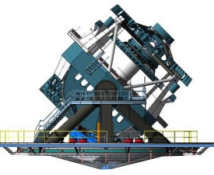
Rolling cadence has impact on discovery of NEO and MBAs. Have not tested MBCs yet.



	NEO H=22	MBA H=21.25	Trojan H=18	TNO H=7.5
baseline_1exp_pairsame_10yrs	67.3	59.2	57.5	57.7
roll_mod2_sdf0.05mixed_10yrs	65.3	57.0	58.2	56.9
roll_mod2_sdf0.20mixed_10yrs	65.3	57.1	57.3	56.7
roll_mod3_sdf0.05mixed_10yrs	64.3	55.0	58.5	56.5
roll_mod3_sdf0.20mixed_10yrs	64.9	55.9	58.7	56.8

	NEO H=22	MBA H=21.25	Trojan H=18	TNO H=7.5
baseline_1exp_pairsame_10yrs	67.3	59.2	57.5	57.7
roll_mod2_sdf0.05mixed_10yrs	65.3	57.0	58.2	56.9
roll_mod2_sdf0.20mixed_10yrs	65.3	57.1	57.3	56.7
roll_mod3_sdf0.05mixed_10yrs	64.3	55.0	58.5	56.5
roll_mod3_sdf0.20mixed_10yrs	64.9	55.9	58.7	56.8

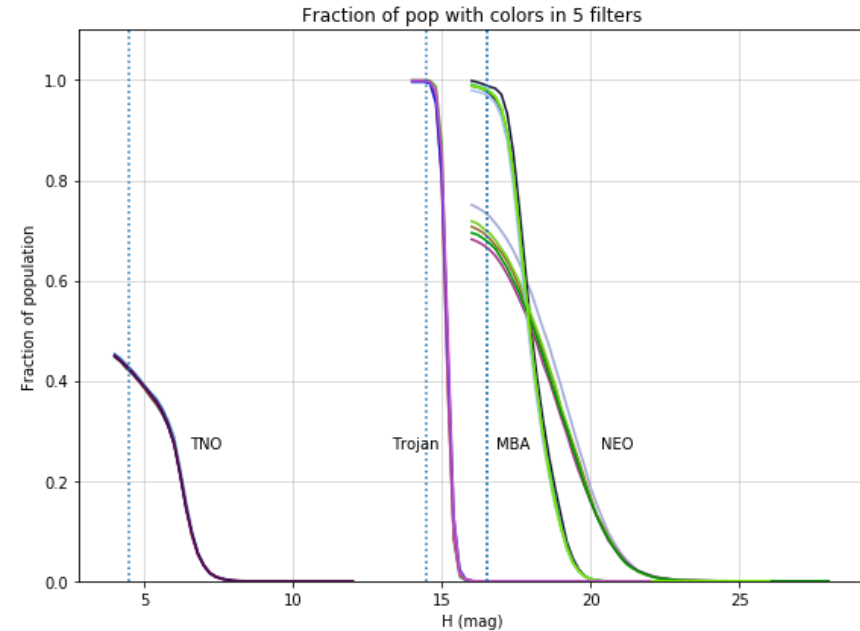
~2-3% losses for NEOs and MBAs, slightly better with higher background visits



# Rolling cadence FBS 1.2 runs



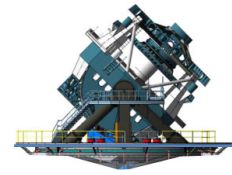
- 'rolling cadence variations'



NEO H=16.5   MBA H=16.5   Trojan H=14.5   TNO H=4.5

	NEO H=16.5	MBA H=16.5	Trojan H=14.5	TNO H=4.5
baseline_1exp_pairsame_10yrs	73.3	98.9	100.0	42.6
roll_mod2_sdf0.05mixed_10yrs	68.9	97.7	99.8	41.9
roll_mod2_sdf0.20mixed_10yrs	69.9	98.1	99.8	42.2
roll_mod3_sdf0.05mixed_10yrs	66.6	97.0	99.5	42.1
roll_mod3_sdf0.20mixed_10yrs	67.9	97.9	99.8	42.3

~4-7% fewer NEOs  
obtaining  
measurement of  
grizy colors



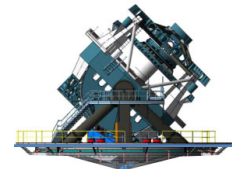
- Rolling cadence
- Motivated by desire to increase cadence for WFD observations (better discovery for transients)
  - Rolling cadence does better for SN discovery, but doing pairs in mixed filters is largest improvement
- Has some negative impact on discovery and characterization of inner solar system (NEO, MBA) objects
  - Check wider range of rolling cadence runs
  - Run simulations with higher background rate?
- Likely to need full-sky coverage each year for difference imaging templates & calibration
- Likely we're missing some metrics sensitive to rolling cadence variations



# What's next? (FBS 1.3 and beyond)



- Additions
  - Fix bug in depth calculation (1x30 vs 2x15)
  - Improve DD sequences (DESC) and add AGN DD sequences
  - Add mini-survey variations (N, S, GP) to evaluate range of impacts on time requirements
  - Run footprints with WFD held at 90%, as well as held at 825 visits/pointing
- BUT need improvements in metrics as well.
  - Need to push metric development (work with Fed)
  - Solar system metrics need addition of MBC and resonant TNO populations (more sensitive)
  - Need more samples of transients (and requirements)
  - Footprint metric (galaxy counts?)



## What's next? (FBS 1.3 and beyond)



- Beyond 1.3:
  - Bright planet (and satellite) avoidance
  - Add more more sophisticated sequences for WFD (specifications on filters for next-night observations)
  - Tackle remaining queue
- Run releases every other month (Sep, Nov, Jan.)
- Write ups and respond to what we're learning

### Still in the queue

- AGN DDF
- Akari and WFIRST/Euclid DDF experiments
- Bulge and low galactic latitude variations
- LMC/SMC mini-surveys
- Twilight NEO survey
- Twilight DCR
- Mini-surveys in the North
- Season extension (not super well defined)
- Anti-alias timing (is it really a problem?)